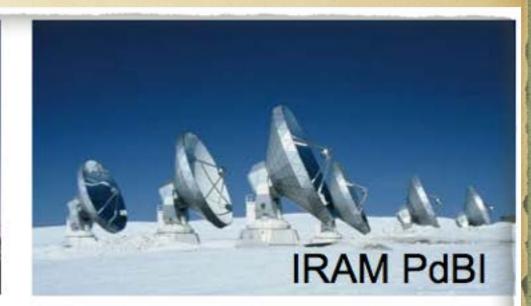
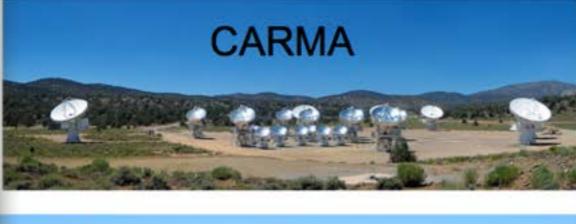
Introduction of Synthesis Imaging in Radio Astronomy

Hui Shi

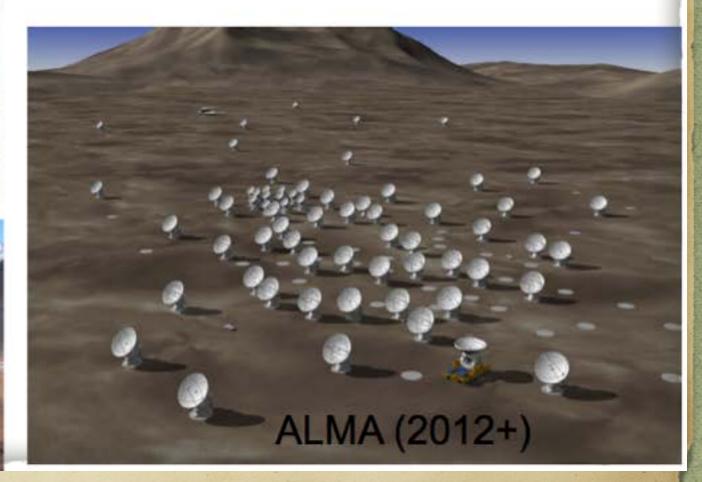




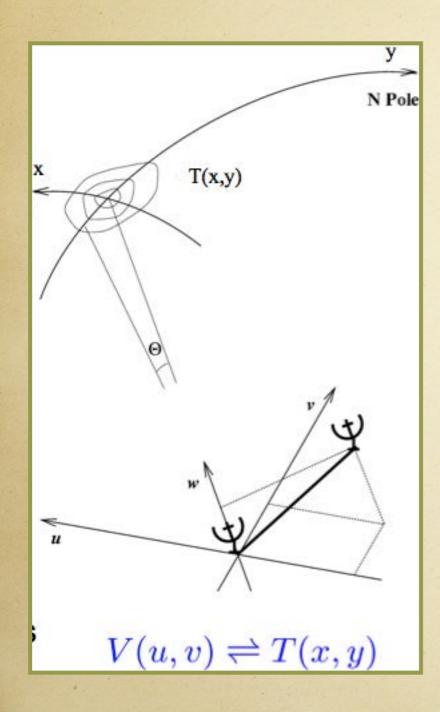








Visibility & Sky Brightness



$$V(u,v) = \int \int T(x,y)e^{2\pi i(ux+vy)}dxdy$$

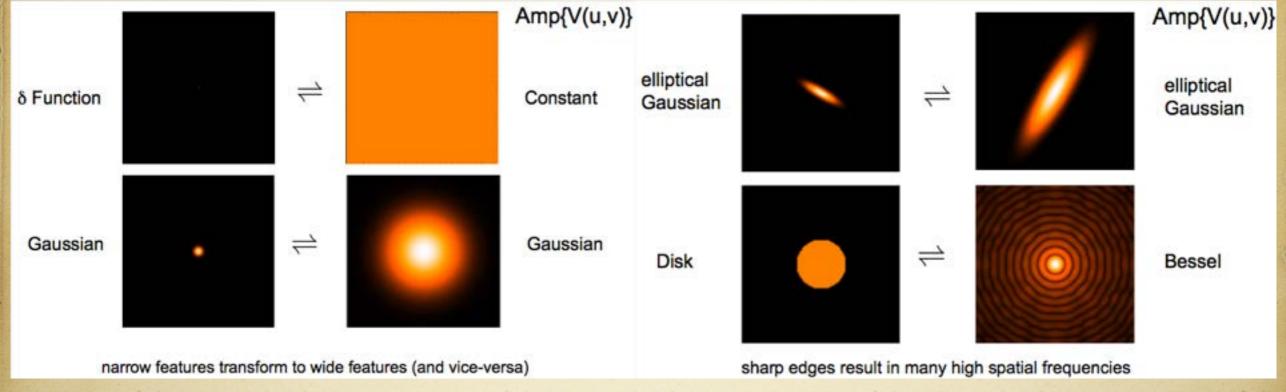
$$T(x,y) = \int \int V(u,v)e^{-2\pi i(ux+vy)}dudv$$

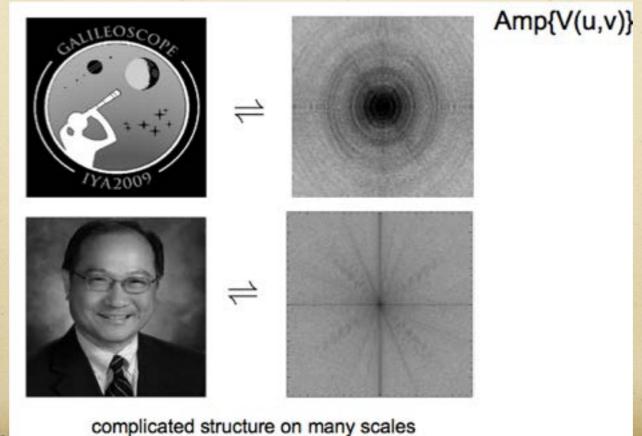
Complex Numbers:

(real, imaginary) or (amplitude, phase)

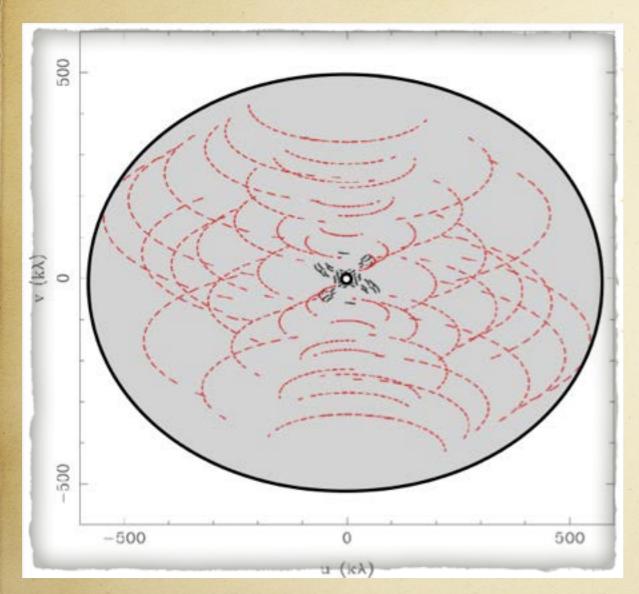
- > Amplitude: "how much" of a certain frequency component
- Phase: "where" this component is located

Fourier Transform Pairs





(u, v) plane sampling



Limited by:

- number of telescopes
- earth-sky geometry

Visibilities: (amplitude, phase)

V(u,v)

Sample:

$$B(u,v) = \sum_{k} (u_k, v_k)$$

Inverse Fourier transform:

$$T^D(x,y) = FT^{-1}\{B(u,v)\times V(u,v)\}$$

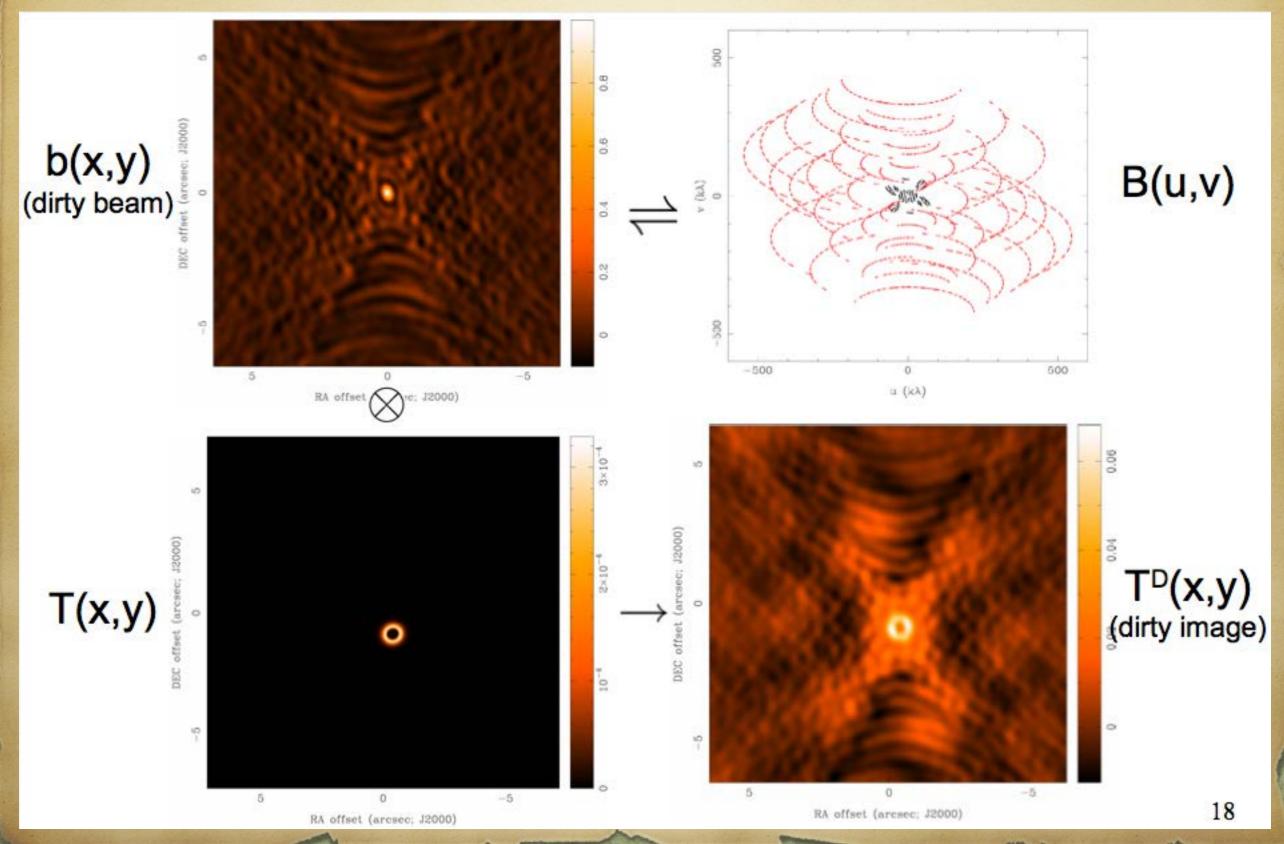
Dirty Beam:

$$b(x,y) = FT^{-1}\{B(u,v)\}\$$

Dirty Image:

$$T^D(x,y) = b(x,y) \otimes T(x,y)$$

Dirty Beam & Dirty Image



Visibilities

V(u,v)

 $B(u,v) = \sum_{k} (u_k, v_k)$

Synthesis Calibrations

Fourier transform

 $b(x,y) = FT^{-1}\{B(u,v)\}$

Dirty Beam -> Clean Beam Dirty image -> Model

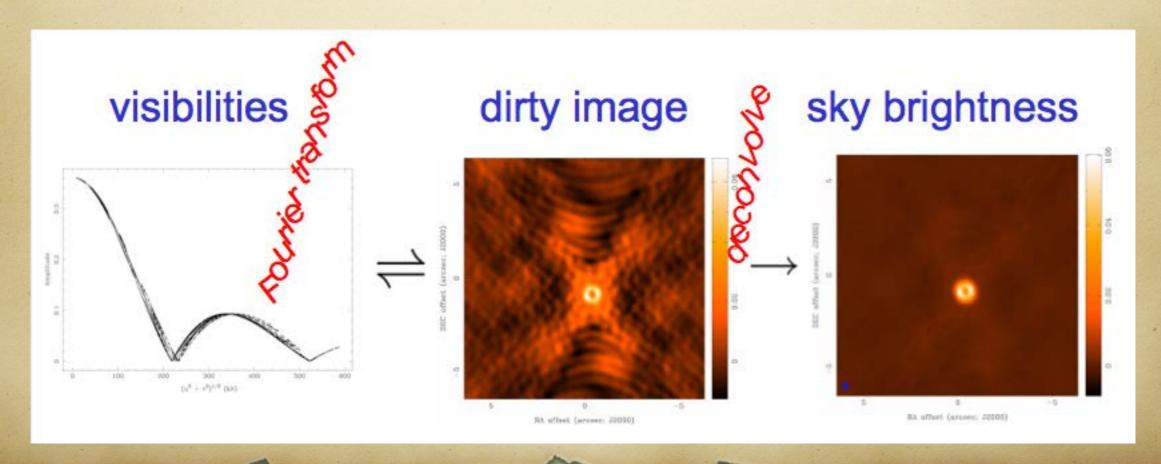
 $T^D(x,y) = b(x,y) \otimes T(x,y)$

Deconvolve

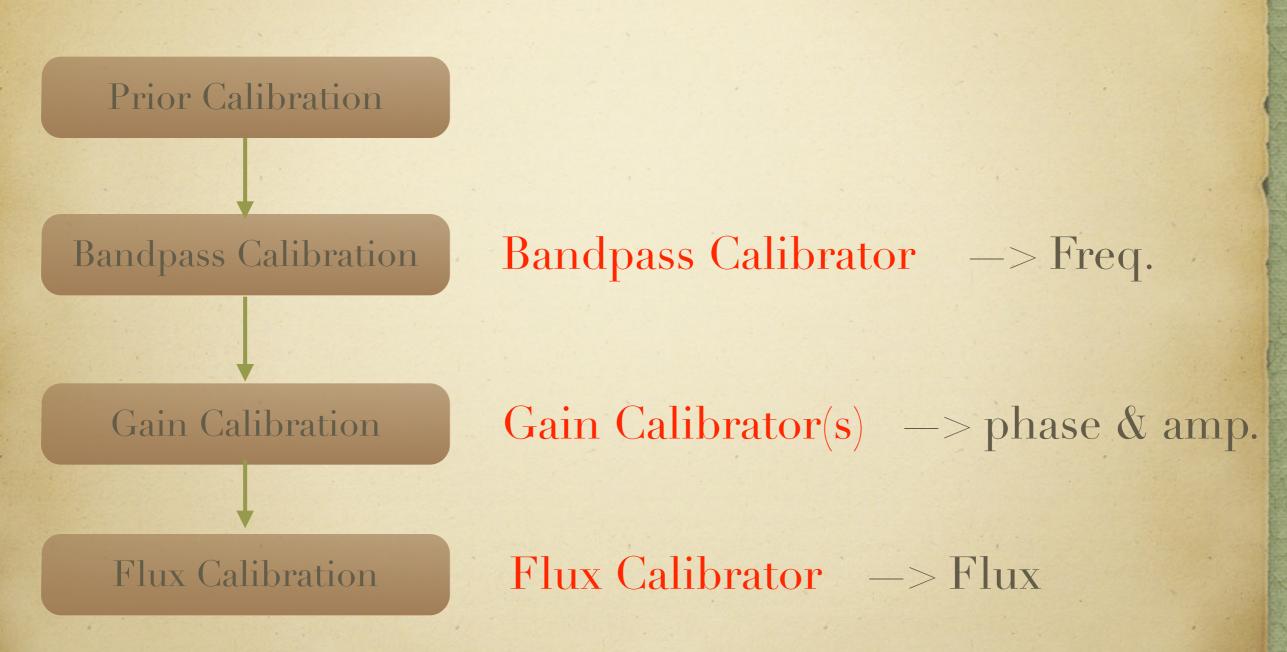
Clean

Clean Image

Analysis

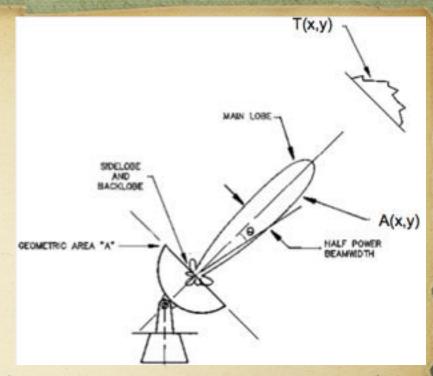


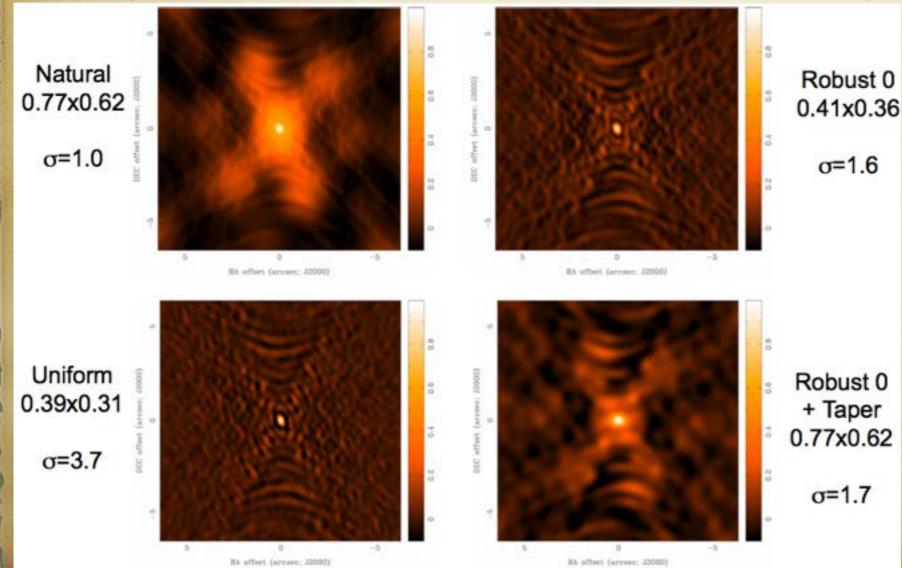
Synthesis Calibrations (to visibility data)

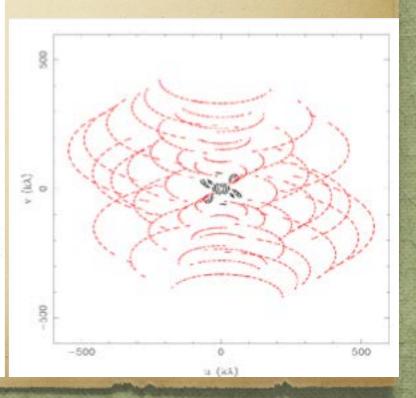


Beam

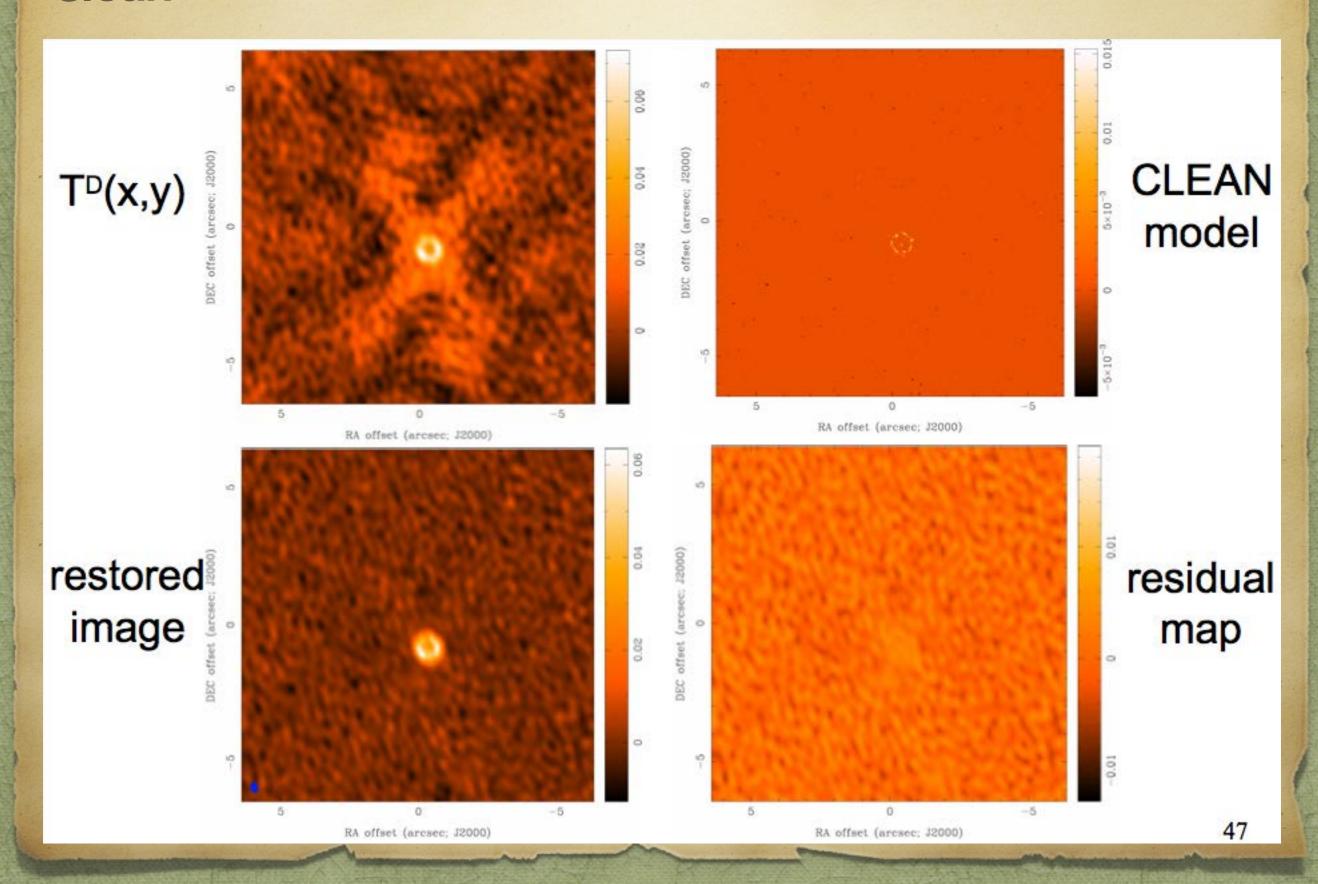
- > Primary Beam —> Field of View (for one single pointing)
- Dirty Beam
 - ⇒ Weighting —> Resolution







Clean



CASA Tutorials

- > Basic Commands
- Data Structure: Measurement Set (MS)
- Data Reduction
 - > Examination
 - > Flagging
 - > Prior Calibration
 - > Main Calibration
 - Clean
- > Image Analysis

> startup

- casapy / casapy nologer
- casalogger
- casapy-YYYYMMDD-HHMMSS.log
- casapy logfile otherfile.log
- > columns
 - > Time: generated time
 - > Priority: priority level
 - casalog.filter('...')
 - > Origin: Task::Tool::Method
 - > Message: actual text
- > features
 - > Search
 - > Filter
 - > View
 - Insert message
 - > Copy
 - Open

> exit

- quit (ask 'y or n')
- > exit
- > CTRL-D
- help
 - > task?
 - %pdoc task (cleanest)
 - > help task
 - > help 'task'
 - > help tool.methd
 - > help
- history
- > Terminal commends
 - > !rm -r mydata.ms
 - rm -r mydata.ms
 - os.system('rm -r mydata.ms')
 - > rmtables('mydata.ms')

Basic Commands

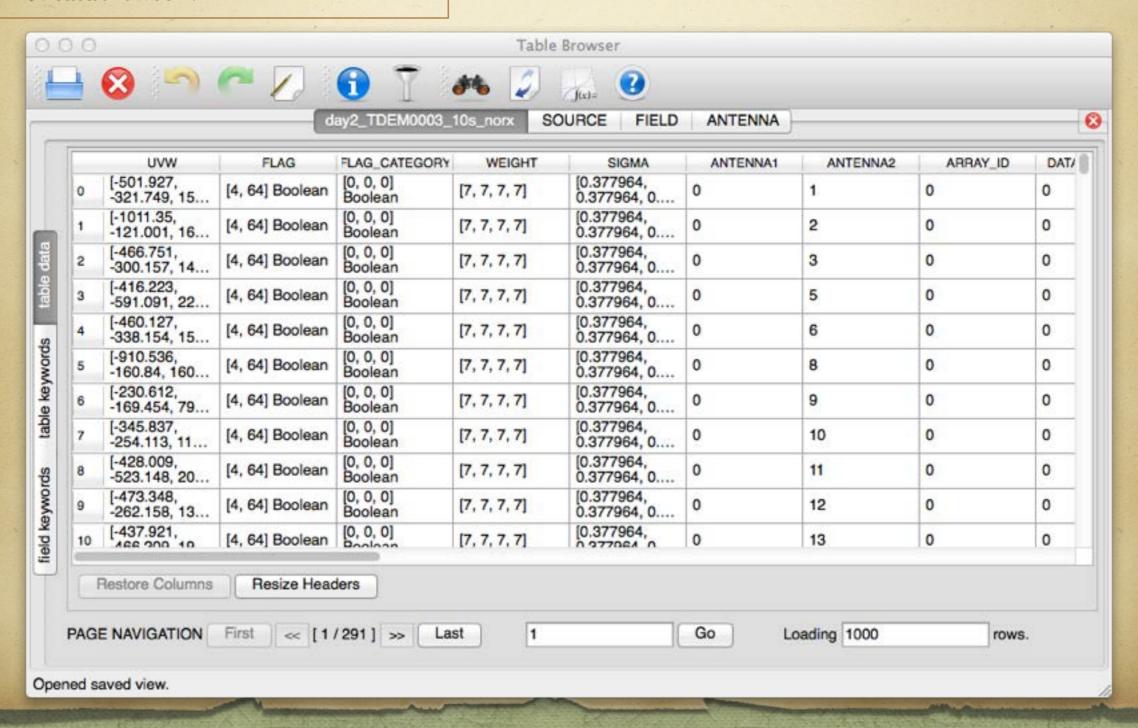
- > View the task parameters
 - > inp task
 - > tget task & inp
- > execute tasks
 - > task
 - op go
 - > task(...)
 - > execfile('myscript.py')
 - > execfile 'myscript.py'
 - > run myscript.py
 - run task.last
 - \$: casapy -c 'myscript.py'

Data structure: Measurement Set (MS)

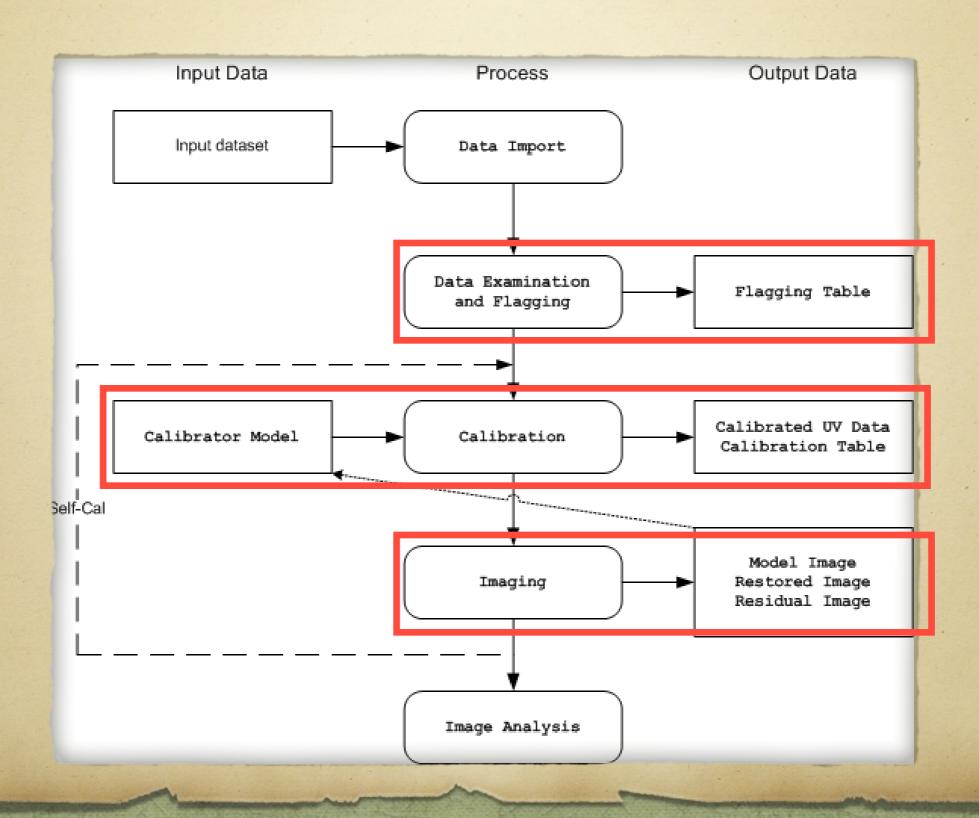
> browserdata

⇒ \$: casabrowser.

Measurement Set (MS)



Data Reduction (Flow Chart)



Load the data to CASA

- > importuvfits (UVFITS, visibility data)
- > importvla (VLA data)
- > importasdm (ALMA data in ASDM format)
- > importevla (JVLA/EVLA data in SDM format)
- > importfits (FITS, image)

Data Examination

- > listobs (summarize)
- plotants
- plotms (visibility data)
- > viewer (image)
- > plotcal (calibration solutions)
- vishead

1. Import

2. Examination

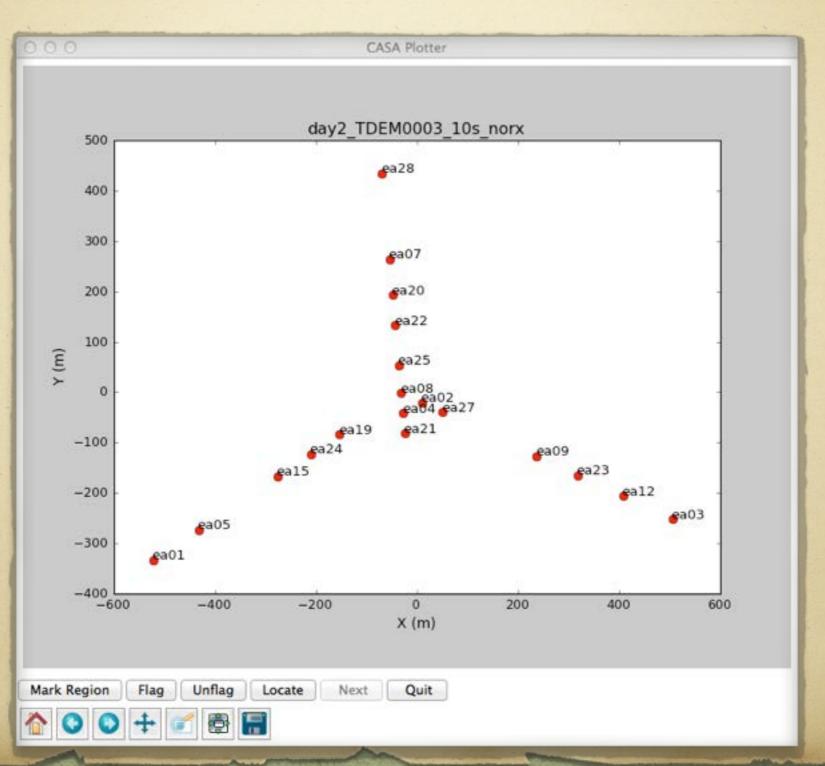
- TO 1000										
Fields	s: 4									
ID	Code Name		R	A	Decl	Epoch	Src	Id		nRows
2	D J0954+	1743	0	9:54:56.82362	26 +17.43.31.2	2243 J2000	2			65326
3	NONE IRC+10	216	09	9:47:57.38200	00 +13.16.40.6	5999 J2000	3		2	08242
5	F J1229+	0203	12	2:29:06.69972	29 +02.03.08.5	9820 J2000	5			10836
7	E J1331+	3030	1:	3:31:08.28798	84 +30.30.32.9	5886 J2000	7			5814
Spects	ral Windows:	(2 uniqu	e speci	tral windows	and 1 unique	polarization	set	ups)		
Spw	ID Name	#Chans	Frame	Ch0(MHz)	ChanWid(kHz)	TotBW(kHz)	Cor	rs		
0	Subband:0	64	TOPO	36387.229	125.000	8000.0	RR	RL	LR	LL
1	Subband:0	64	TOPO	36304.542	125.000	8000.0	RR	RL	LR	LL
Source	es: 10									
ID	Name		SpwId I	RestFreq(MHz)) SysVel(km/s)				
0	J1008+0730		0 (0.03639232	-0.026					
0	J1008+0730		1 (0.03639232	-0.026					
2	J0954+1743		0 (0.03639232	-0.026					
2	J0954+1743		1 (0.03639232	-0.026					
3	IRC+10216		0 (0.03639232	-0.026					
3	IRC+10216		1 (0.03639232	-0.026					
5	J1229+0203		0 (0.03639232	-0.026					
5	J1229+0203		1 (0.03639232	-0.026					
7	J1331+3030		0 (0.03639232	-0.026					
7	J1331+3030		1 (0.03639232	-0.026					

Examination: listobs

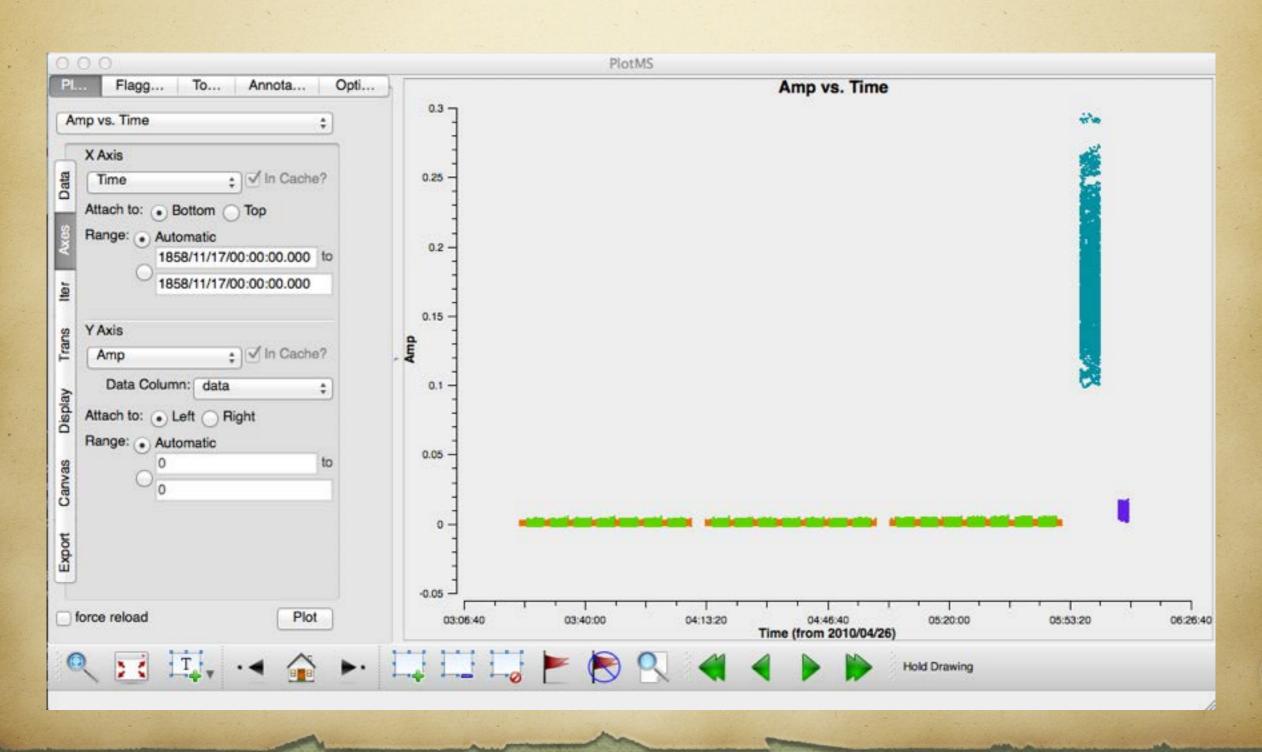
Observer: Mark J. Mark Claussen Project: T.B.D. Observation: EVLA Data records: 290218 Total integration time = 10016 seconds Observed from 26-Apr-2010/03:21:56.0 to 26-Apr-2010/06:08:52.0 (UTC) ObservationID = 0 ArrayID = 0 Scan FldId FieldName Average Interval(s) ScanIntent Date Timerange (UTC) nRows SpwIds 26-Apr-2010/03:21:51.0 - 03:23:21.0 2 J0954+1743 2720 [0, 1] [10, 10] 9918 [0, 1] 03:23:39.0 - 03:28:25.0 3 IRC+10216 [10, 10] 03:28:38.0 - 03:29:54.0 2 J0954+1743 2700 [0, 1] [10, 10] 03:30:08.0 - 03:34:53.5 3 IRC+10216 9918 [0, 1] [10, 10] 03:35:07.0 - 03:36:23.0 2 J0954+1743 2736 [0, 1] [10, 10] 03:36:37.0 - 03:41:22.5 10 3 IRC+10216 9918 [0, 1] [10, 10] 2700 [0, 1] [10, 10] 03:41:35.4 - 03:42:51.5 11 2 J0954+1743 03:43:05.8 - 03:47:51.5 3 IRC+10216 9918 [0, 1] [10, 10] 03:48:04.4 - 03:49:20.5 2 J0954+1743 2700 [0, 1] [10, 10] 03:49:34.9 - 03:54:20.5 3 IRC+10216 9918 [0, 1] [10, 10] 03-54-33 0 - 03-55-49 6 2 .10954+1743 2496 10. 11 110. 101

ID Name Station Diam. Long. Lat. Offset from array center (m) ITRF Geo	ocentric coordinates (m)
ID Name Station Diam. Long. Lat. Offset from array center (m) ITRF Geo	
East North Elevation x	y z
0 ea01 W09 25.0 m -107.37.25.2 +33.53.51.0 -521.9407 -332.7782 -1.1977 -1601710.017000	-5042006.928200 3554602.355600
1 ea02 E02 25.0 m -107.37.04.4 +33.54.01.1 9.8247 -20.4292 -2.7808 -1601150.059500	-5042000.619800 3554860.729400
2 ea03 E09 25.0 m -107.36.45.1 +33.53.53.6 506.0591 -251.8666 -3.5832 -1600715.948000	-5042273.187000 3554668.184500
3 ea04 W01 25.0 m -107.37.05.9 +33.54.00.5 -27.3562 -41.3030 -2.7418 -1601189.030140	-5042000.493300 3554843.425700
4 ea05 W08 25.0 m -107.37.21.6 +33.53.53.0 -432.1158 -272.1493 -1.5032 -1601614.091000	-5042001.655700 3554652.509300
5 ea07 NO6 25.0 m -107.37.06.9 +33.54.10.3 -54.0667 263.8720 -4.2292 -1601162.593200	-5041829.000000 3555095.890500

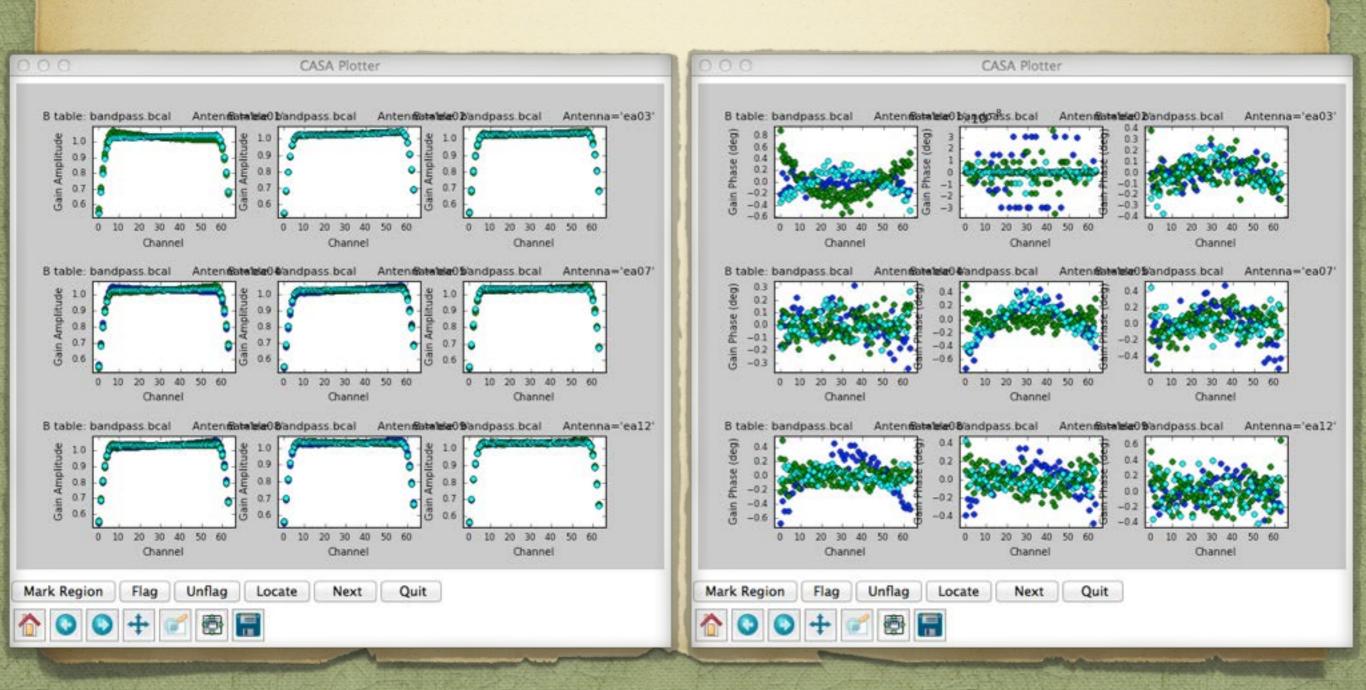
Examination: plotants



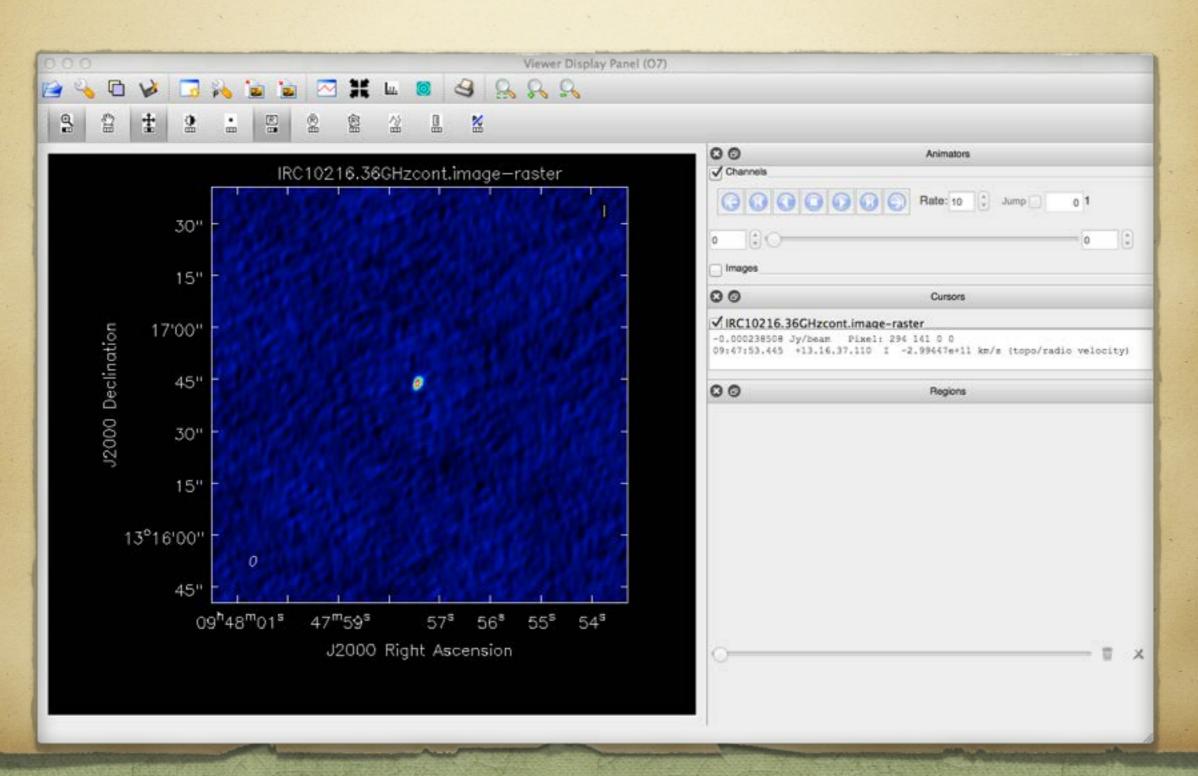
Examination: plotms (Visibility Data)



Examination: plotcal (calibration solutions)



Examination: viewer (Image Data)

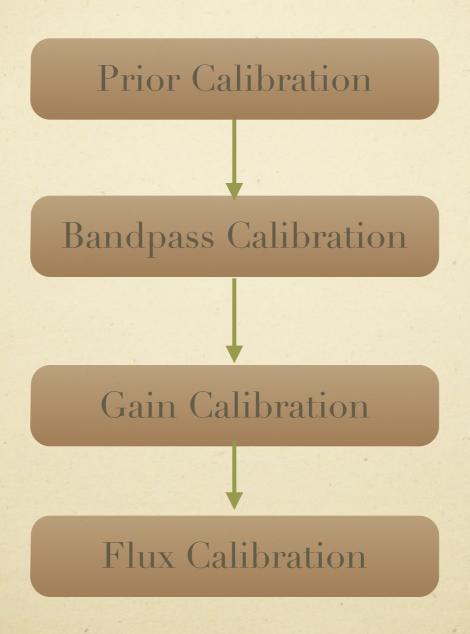


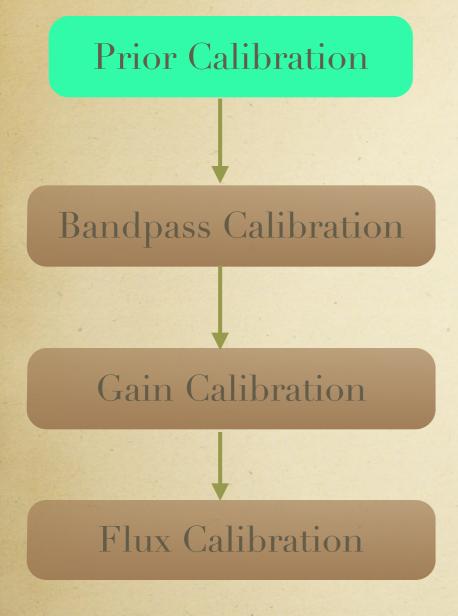
Data flagging

- > flagdata
 - > mode='manual'
 - > mode='list'
 - > mode='unflag'
 - > mode='summary'
 - > mode='clip' ('quack', 'shadow', 'elevation', 'tfcrop', 'rflag', 'extend')
- flagmanager (MS)
 - > mode='list'
 - > mode='save'
 - > mode='restore'
 - > mode='delete'
 - > mode='rename'
- > flagemd
 - action='apply' ('unapply','list', 'plot','clear','extract')
- > Interactive flagging (backup before do that)
- > Auto-RFI flagging
 - ⇒ flagdata: mode='rflag'

3. Flagging

4. Synthesis Calibrations





System Temperature & Switched-Power Correction

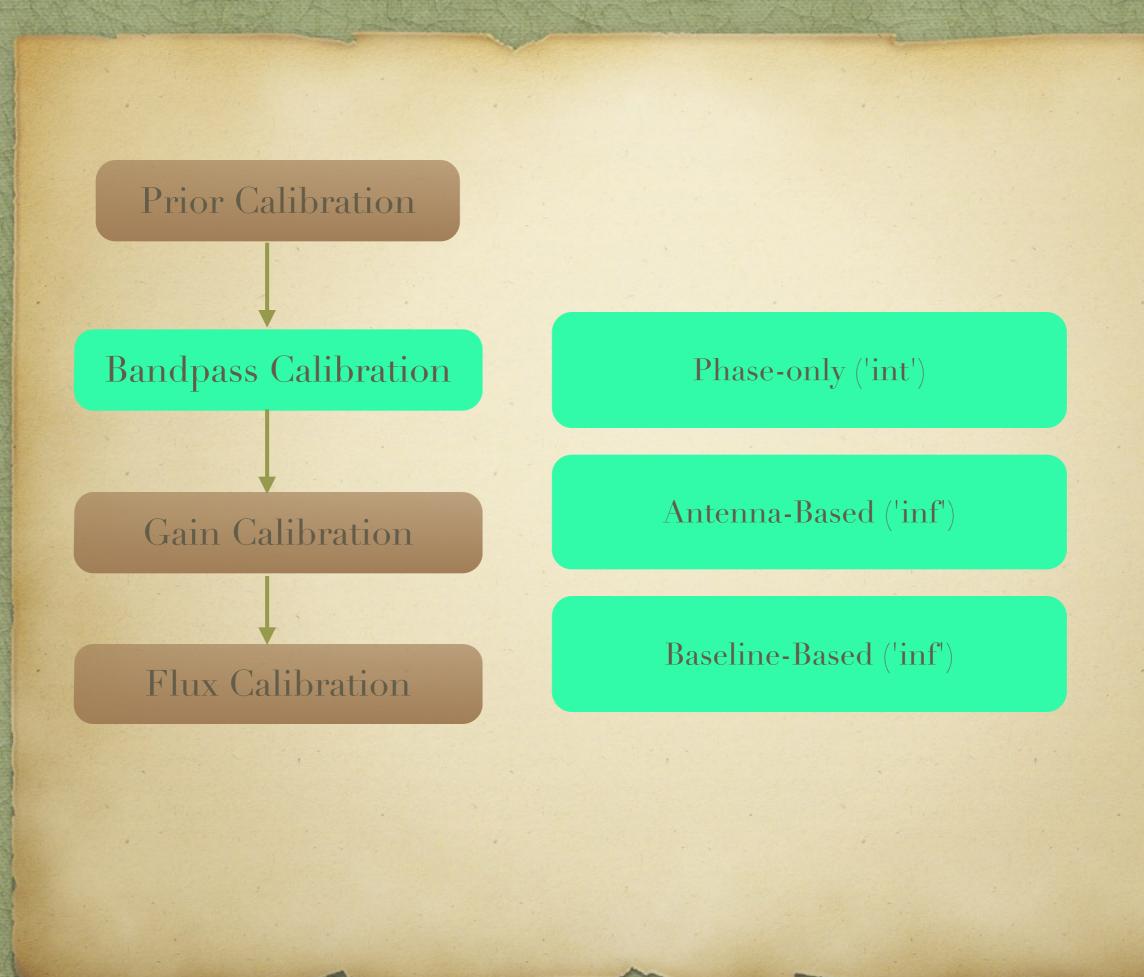
Flux Density Scale (model)

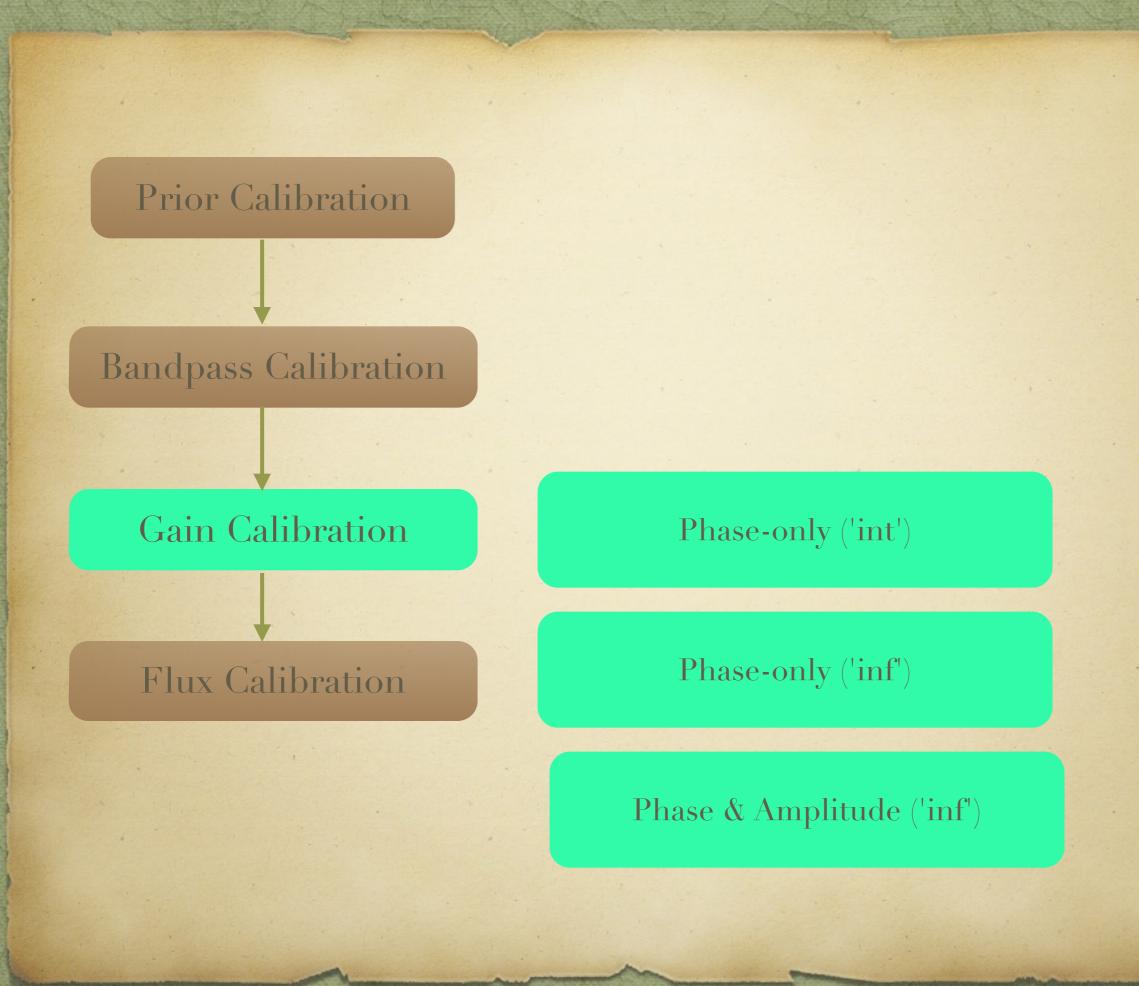
Antenna Gain-Elevation Curve Calibration

Atmospheric Optical Depth Correction

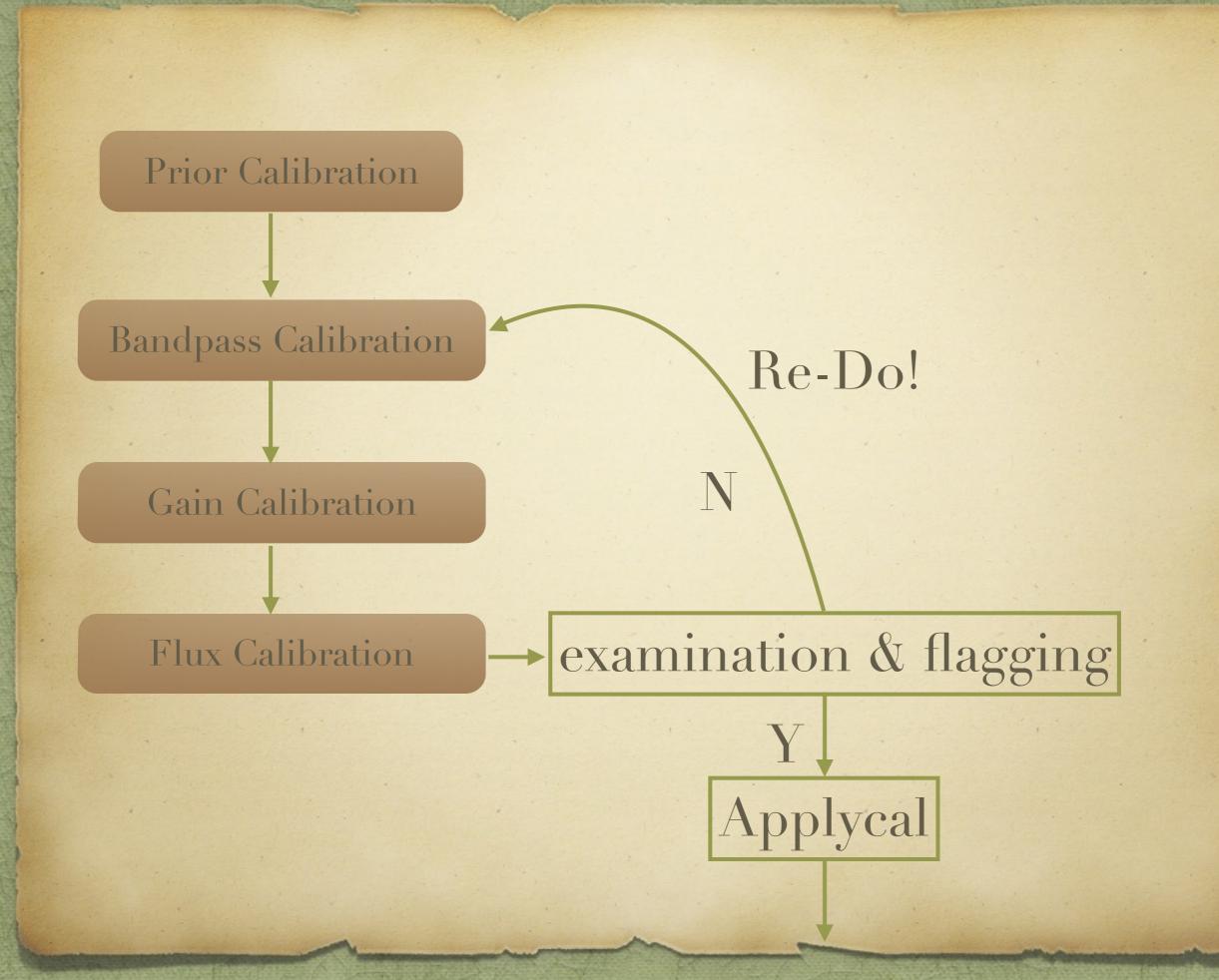
Antenna-based Delay (sbd, mbd)

Antenna position offset Correction

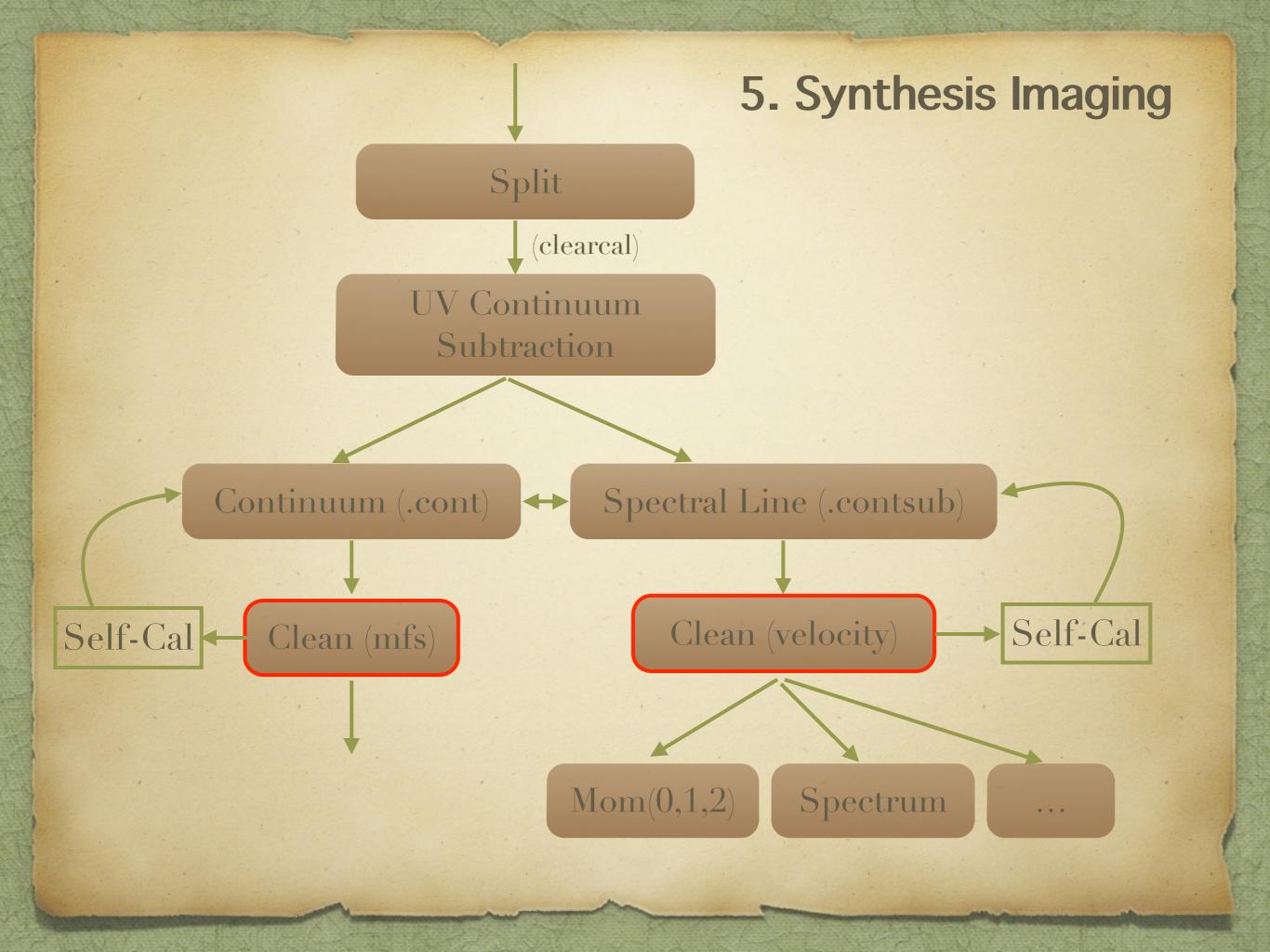




Prior Calibration Bandpass Calibration Gain Calibration Flux Calibration



- > Important Tips:
 - > check & check
 - > take a note
 - > backup



- > cell: pixel size
- > imsize: 10*2^n (n=1,2,3,4,5,6...)
- > weighting
 - \Rightarrow natural $w_i = \omega_i = \frac{1}{\sigma_k^2}$
 - \Rightarrow uniform & superuniform $w_i = \frac{\omega_i}{W_k}$.
 - radial $w_i = \omega_i \cdot \sqrt{u_i^2 + v_i^2}$.
 - > briggs $w_i = \frac{\omega_i}{1 + W_k f^2}$ $f^2 = \frac{(5 * 10^{-R})^2}{\sum_k W_k^2}$
 - > robust: -2(uniform) 2.0(natural)
- > mode
 - > mss
 - > velocity/channel/frequency
- > psfmode
 - > 'clark' algorithm (default)
 - > 'hogbom' algorithm
- imagermode
 - csclean
 - mosaic

- > threshold: rms
- > niter
- > gain: 0.1
- multiscale
- > mask
- interactive

Clean

- > imagename.image > restored image
- imagename.flux > effective response
- > imagename.flux.pbcoverage > PB coverage (ftmachine='mosaic')
- imagename.model > model image (Jy/pix)
- ⇒ imagename.psf > synthesized (dirty) beam

- > Image analysis & viewing (image cube)
 - > moment (immoment)
 - > spectral line fitting (specfit)
 - > P-V diagram (impv)
 - image-component fitting (imfit)
 - image convolution (imsmooth)
 - continumm subtraction (imcontsub)
 - > (imhead)
 - statistics (imstat)
 - > math (immath)
 - > regride (imregrid)
 - > redefine the velocity frame (imreframe)
 - > ...

- > Self-calibration
- Mosaic
 - task 'clean'
 - > imagermode='mosaic'
 - ftmachine='ft' -> slow (weighted combination of images)
 - ftmachine='mosaic'(default)
 - -> fast (gridded onto a single uv-plane -> single output image)
 - —> needed primary beam correction
 - -> twice as big as necessary to encompass the mosaic
 - > minpb
 - > phasecenter
 - > Toolkit 'MEM'

Thanks!